

# Supplemental: An Approximate Mie Scattering Function for Fog and Cloud Rendering

Johannes Jendersie\*  
 jjendersie@nvidia.com  
 NVIDIA  
 Berlin, Germany

Eugene d'Eon\*  
 edeon@nvidia.com  
 NVIDIA Research  
 Wellington, New Zealand

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## 1 OUTLINE

In this supplemental we provide a number of additional results:

- section 2 provides full details of the importance sampling of our fog phase function, which includes new sampling methods for the Draine and Cornette-Shanks phase functions.
- section 3 provides additional parametric fitting functions for particle diameters spanning  $0 < d < 50\mu\text{m}$ , together with plots illustrating the accuracy of these approximations.

## 2 IMPORTANCE SAMPLING

Our fog phase function is the sum of two non-negative lobes,

$$\phi_{\text{fog}}(\theta) = (1 - w_D) \cdot \phi_{\text{HG}}(\theta) + w_D \cdot \phi_{\alpha, g_D}(\theta). \quad (1)$$

which we sample in two steps. We first select one of the two lobes using mixture weight  $w_D$  and then sample the selected lobe using the procedures listed below. Since both lobes in our model are themselves normalized phase functions, the Draine lobe is sampled when  $\xi_s < w_D$ . In practice, the random number  $\xi_s \in [0, 1)$  used for lobe selection can be rescaled and used again during lobe sampling.

### 2.1 Henyey-Greenstein

Using notation  $u = \cos \theta$ , the HG phase function is [Henyey and Greenstein 1941]

$$\phi_{\text{HG}}(u) = \frac{1}{4\pi} \frac{1 - g^2}{(1 + g^2 - 2gu)^{3/2}}, \quad (2)$$

and can be sampled via [Toublanc 1996; Witt 1977]

$$\begin{aligned} t &= \frac{1 - g^2}{1 - g + 2g\xi} \\ u &= \frac{1 + g^2 - t}{2g} \end{aligned} \quad (3)$$

\*Both authors contributed equally to this research.

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for a random number  $\xi \in [0, 1)$ . If  $g = 0$ , or very small, it is better to use isotropic sampling as a fallback, simplifying to

$$u = 2\xi - 1. \quad (4)$$

### 2.2 Draine

Draine's phase function is a two-parameter distribution,

$$\phi_{\alpha, g}(\theta) = \frac{1}{4\pi} \frac{1 - g^2}{(1 + g^2 - 2g \cos \theta)^{3/2}} \frac{1 + \alpha \cos^2 \theta}{1 + \alpha(1 + 2g^2)/3}, \quad (5)$$

which generalizes the Cornette-Shanks (CS) phase function [Cornette and Shanks 1992]. The original CS model is recovered when  $\alpha = 1$ . It has been claimed that CS cannot be sampled using CDF inversion [Toublanc 1996]. On the contrary, we found (using Mathematica) that the CDF for Draine cosines  $u = \cos \theta$  is

$$\xi = \frac{W(u) + 2\alpha g^5 (V(u) + u) + (\alpha + 3)g^3 V(u) - 2\alpha (V(u) - 1) - 2\alpha g u}{2g^3 (2\alpha g^2 + \alpha + 3) V(u)}$$

$$W(u) = -2\alpha g^6 + g^4 (\alpha (u^2 - 2) - 3) - g^2 (\alpha (V(u) + u^2 - 2) + 3 (V(u) - 1))$$

where  $V(u) = \sqrt{g^2 - 2gu + 1}$  and  $\xi$  is a random number drawn uniformly from  $[0, 1)$ . After a change of variable  $u \rightarrow (1 + g^2 - v^2)/(2g)$  to remove the square root,  $v$  follows from solution of a quartic equation, which produces  $u$  using the above mapping. After factorization of common terms, we find the following form:

$$\begin{aligned} t_0 &= \alpha - \alpha g^2 \\ t_1 &= \alpha g^4 - \alpha \\ t_2 &= -3 \cdot (4(g^4 - g^2) + t_1 \cdot (1 + g^2)) \\ t_3 &= g \cdot (2\xi - 1) \\ t_4 &= 3g^2 \cdot (1 + t_3) + \alpha \cdot (2 + g^2 \cdot (1 + (1 + 2g^2) \cdot t_3)) \\ t_5 &= t_0 \cdot (t_1 t_2 + t_4^2) + t_1^3 \\ t_6 &= t_0 \cdot 4(g^4 - g^2) \\ t_7 &= \left( t_5 + \sqrt{t_5^2 - t_6^3} \right)^{\frac{1}{3}} \\ t_8 &= 2 \frac{t_1 + t_6/t_7 + t_7}{t_0} \\ t_9 &= \sqrt{6(1 + g^2) + t_8} \\ u &= \frac{g}{2} + \left( \frac{1}{2g} - \frac{1}{8g} \left( \sqrt{6(1 + g^2) - t_8 + 8t_4/(t_0 t_9)} - t_9 \right)^2 \right), \end{aligned} \quad (6)$$

which exactly samples the Draine (and therefore CS) phase function. A Mathematica implementation is provided with the supplemental material for additional validation.

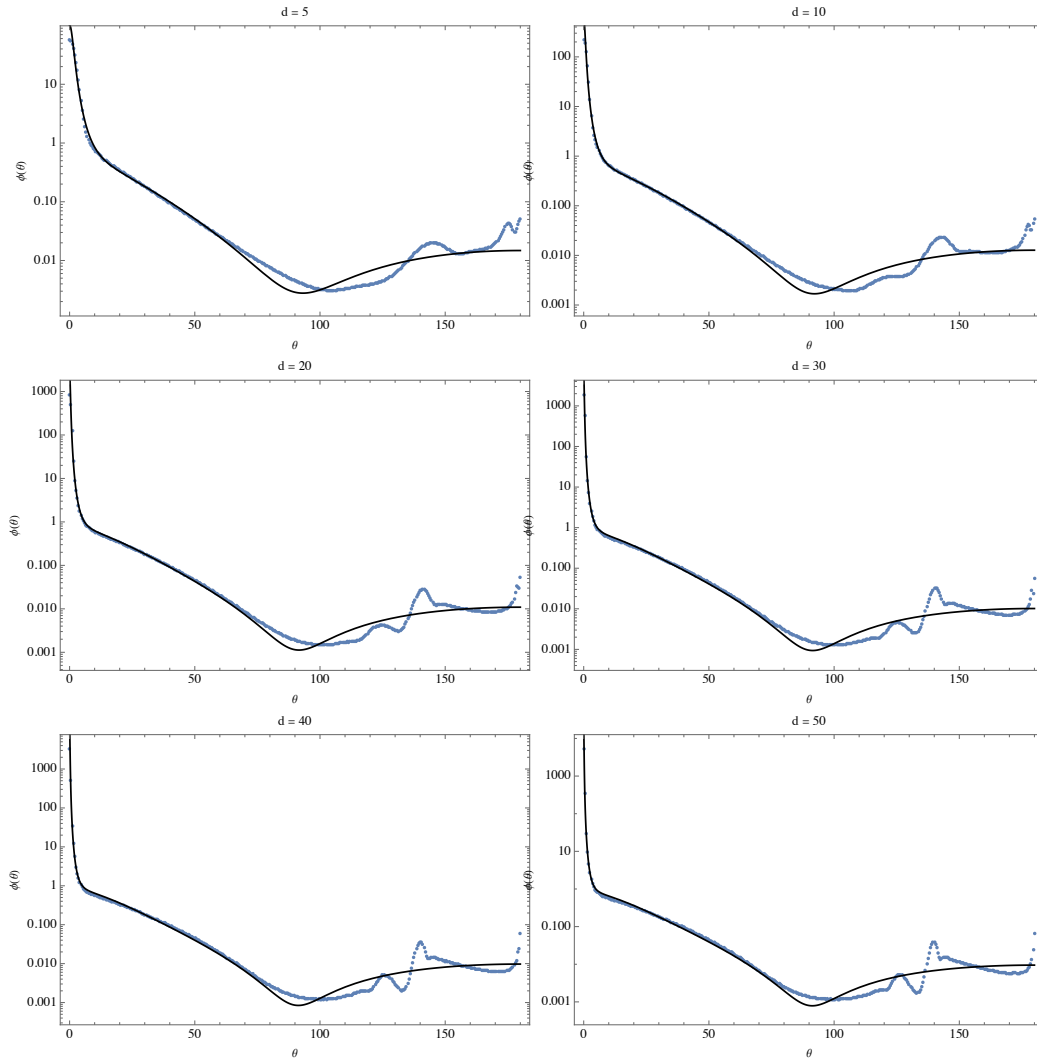


Figure 1: Comparison of our large-particle parametric fit (black) to reference (dots).

### 3 PARAMETRIC FITS

In this section we are comparing plots of our model against the tabulated Mie reference and propose extensions for the range  $d \in [0, 5] \mu\text{m}$ .

#### 3.1 Large particles, Diameter $5 \mu\text{m} \leq d \leq 50 \mu\text{m}$

Figure 1 compares our fitted parametric model for large particles,

$$g_{\text{HG}}(d) = e^{-\frac{0.0990567}{d-1.67154}}, \tag{7}$$

$$g_{\text{D}}(d) = e^{-\frac{2.20679}{d+3.91029} - 0.428934}, \tag{8}$$

$$\alpha(d) = e^{3.62489 - \frac{8.29288}{d+5.52825}}, \tag{9}$$

$$w_{\text{D}}(d) = e^{-\frac{0.599085}{d-0.641583} - 0.665888} \tag{10}$$

(repeated here from the main abstract) to reference data for a variety of particle diameters.

### 3.2 Small particles, Diameter $d \leq 0.1 \mu\text{m}$

For small diameter ( $d < 0.1 \mu\text{m}$ ) we found the following parametric fit to be very accurate:

$$g_{\text{HG}}(d) = 13.8d^2, \quad (11)$$

$$g_{\text{D}}(d) = 1.1456 d \sin(9.29044d), \quad (12)$$

$$\alpha(d) = 250, \quad (13)$$

$$w_{\text{D}}(d) = 0.252977 - 312.983d^{4.3}. \quad (14)$$

Figure 2 shows a comparison of this model to reference data for a variety of small diameters. In the limit of vanishing particle diameter ( $d \rightarrow 0$ ) the phase function approaches Rayleigh scattering, which can be represented in two distinct ways with our model: one that uses the Draine term exclusively:  $w_{\text{D}} = 1, \alpha = 1, g_{\text{HG}} = 0, g_{\text{D}} = 0$ , and one that mixes Draine with HG:  $w_{\text{D}} = 0.25, \alpha = \infty, g_{\text{HG}} = 0, g_{\text{D}} = 0$ . We found that the latter representation served as a more accurate limit for small particle size (when  $d$  is not quite zero), and therefore built the above approximation with  $\alpha = 250$ . This avoids having to special-case the sampling routine for  $\alpha = \infty$ , and results in negligible error.

### 3.3 Mid-range particles, Diameter $0.1 \mu\text{m} < d < 1.5 \mu\text{m}$

Particle diameters  $d$  in the range  $[0.1, 5]$  presented more of a challenge for our two-lobe parametric model. For this range, we split the fitting into two subranges. For particles with diameters  $0.1 < d < 1.5$  we propose

$$g_{\text{HG}}(d) = 0.862 - 0.143 \log^2(d), \quad (15)$$

$$g_{\text{D}}(d) = 0.379685 \cos \left( 1.19692 \cos \left( \frac{(\log(d) - 0.238604)(\log(d) + 1.00667)}{0.507522 - 0.15677 \log(d)} \right) + 1.37932 \log(d) + 0.0625835 \right) + 0.344213, \quad (16)$$

$$\alpha(d) = 250, \quad (17)$$

$$w_{\text{D}}(d) = 0.146209 \cos(3.38707 \log(d) + 2.11193) + 0.316072 + 0.0778917 \log(d). \quad (18)$$

The accuracy of this fit is illustrated in Figure 3.

### 3.4 Mid-range particles, Diameter $1.5 \mu\text{m} \leq d < 5 \mu\text{m}$

For particles with diameters  $1.5 \leq d < 5$  we propose

$$g_{\text{HG}}(d) = 0.0604931 \log(\log(d)) + 0.940256, \quad (19)$$

$$g_{\text{D}}(d) = 0.500411 - \frac{0.081287}{-2 \log(d) + \tan(\log(d)) + 1.27551}, \quad (20)$$

$$\alpha(d) = 7.30354 \log(d) + 6.31675, \quad (21)$$

$$w_{\text{D}}(d) = 0.026914(\log(d) - \cos(5.68947(\log(\log(d)) - 0.0292149))) + 0.376475. \quad (22)$$

The accuracy of this fit is illustrated in Figure 4.

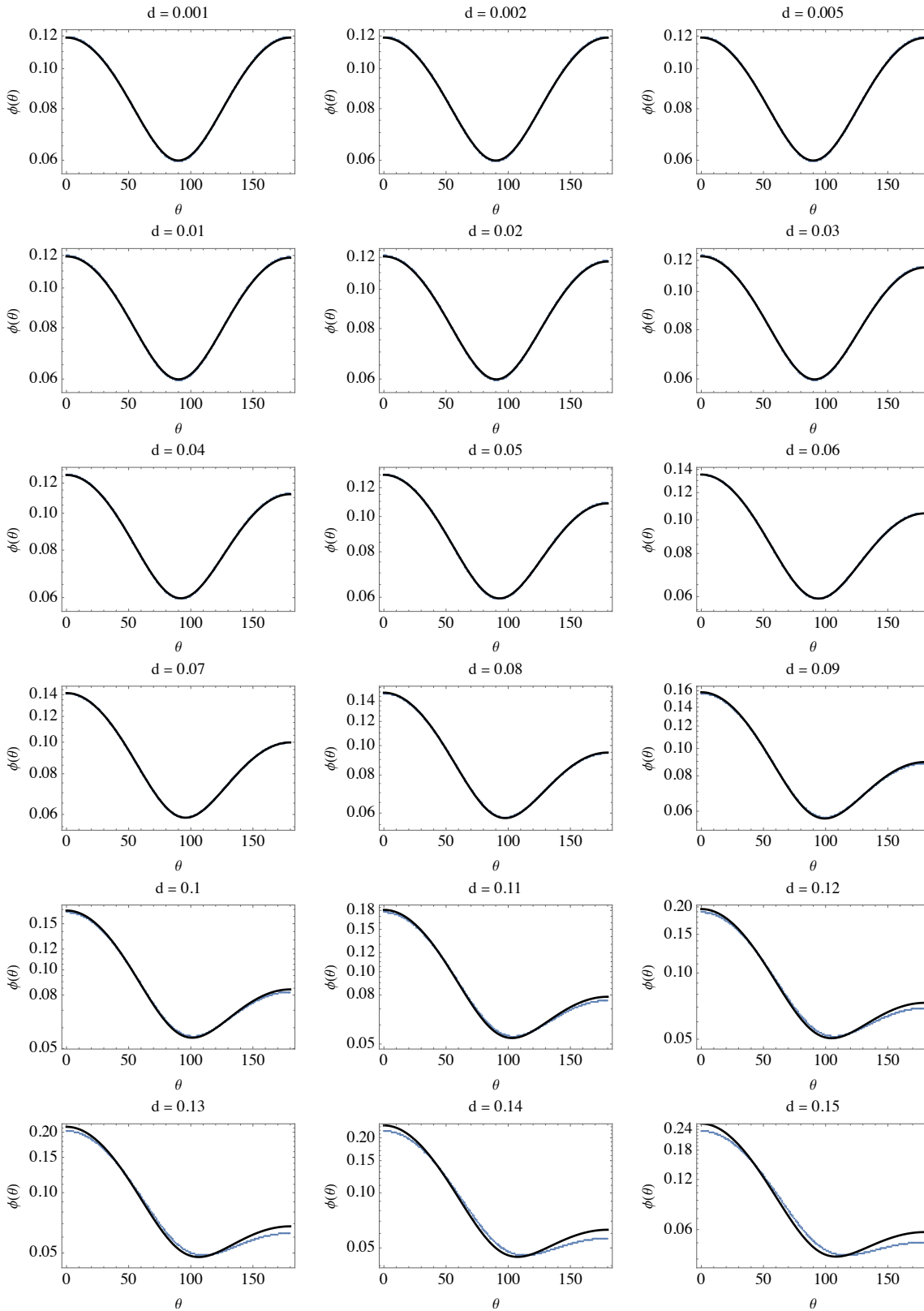


Figure 2: Comparison of our small-particle parametric fit (black, Eq. 11) to reference (dots).

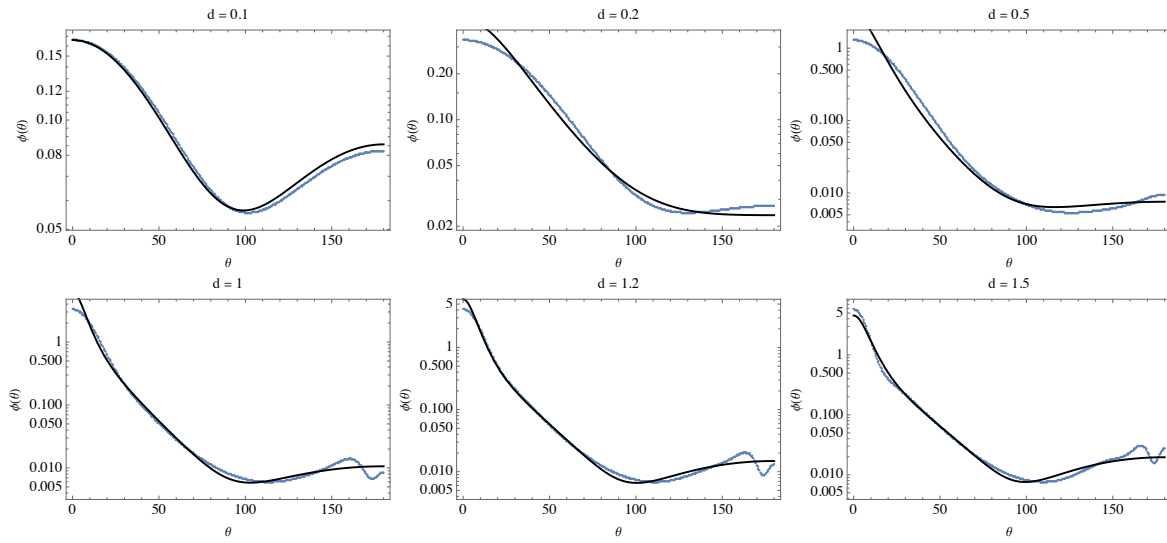


Figure 3: Comparison of our first medium-range-diameter parametric fit (black, Eq. 15) to reference (dots).

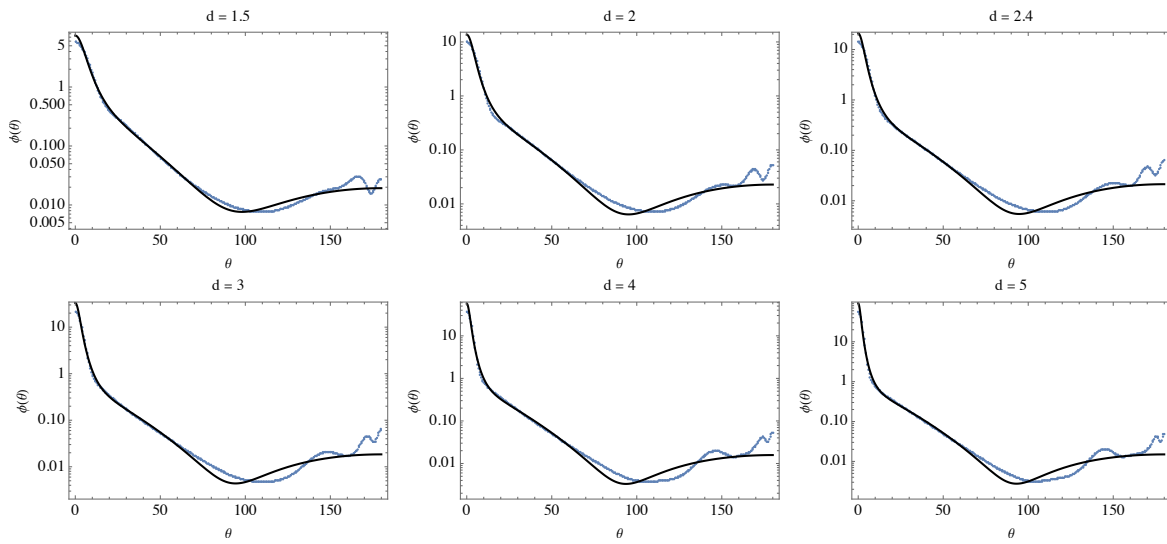


Figure 4: Comparison of our second medium-range-diameter parametric fit (black, Eq. 19) to reference (dots).

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